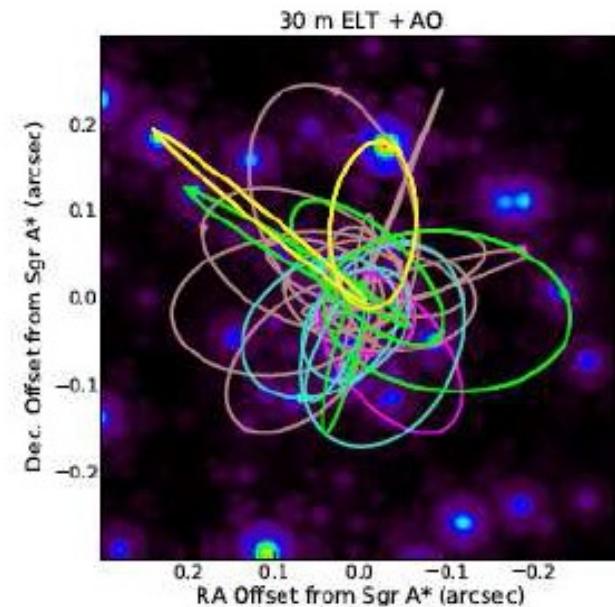
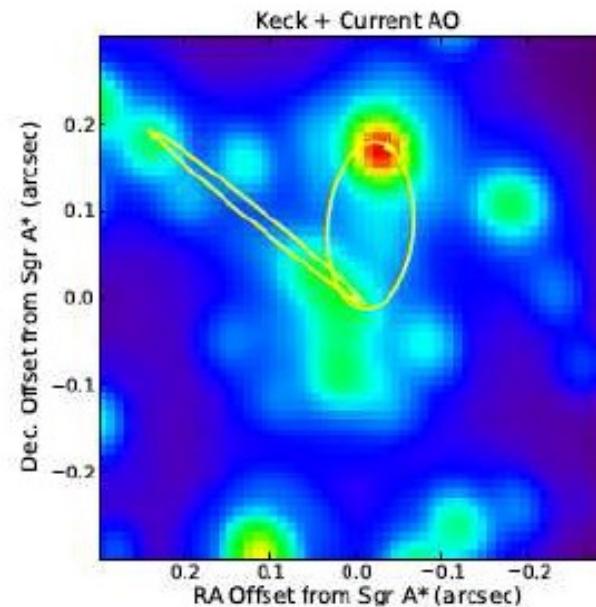
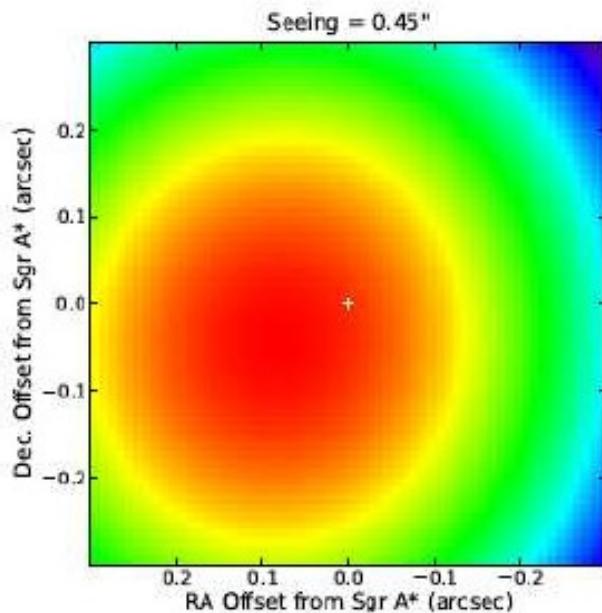


Thirty Meter Telescope - First Light Instrumentation

The Infra-Red Imaging Spectrograph (IRIS)



TMT



30 m 望遠鏡
三十米望远镜
तीस मीटर दूरबीन
Thirty Meter Telescope
Téléscope de Trente Mètres

Caltech

INDIA
TMT



国立天文台
NAOJ
National Astronomical Observatory of Japan



Significant funding provided by the Gordon and Betty Moore Foundation

The Infra-Red Imaging Spectrograph (IRIS)

IRIS is diffraction limited integral field spectrograph and imager designed to be operational at first light using TMT's Adaptive Optics system NFIRAOS

Right: Specification of the different observing modes.

Bottom left: Expected Performances.

Bottom right: Section view of IRIS optical components.

Back: Simulation of TMT performance on the galactic center (courtesy of Ghez et al.).

Capability mode	Spatial sampling (mas)	Field of View (arcsec)	Resolution (λ/d)	Min/Max wavelength (mm)	Bandpass ³	Coronagraph
Imager	4 mas	16.4 x 16.4 Possible expansion to 32.8x32.8	Set by filter	0.84-2.4	37 filters Variety of bandpasses	Planning mask mounted in OIWFS
Slicer IFS	50 mas	4.4 x 2.25	4,000, 8,000	0.84-2.4	20%, 10%	Planning mask mounted in OIWFS
88x45 Spaxels	25 mas	2.2 x 1.125	4,000, 8,000	0.84-2.4	20%, 10%	
Lenslet IFS	9 mas	1.01 x 1.15	4,000	0.84-2.4	5%	Planning mask mounted in OIWFS
112x128 Spaxels	4 mas	0.45 x 0.51	4,000	0.84-2.4	5%	
Lenslet IFS	9 mas	0.144 x 1.15	8,000-10,000	0.84-2.4	20%	Planning mask mounted in OIWFS
16x128 Spaxels	4 mas	0.064 x 0.51	8,000-10,000	0.84-2.4	20%	

Performance category	Value	Comment
Expected Strehl ratio for greater than 50% of sky	J band: 0.41 H band: 0.60 K band: 0.75	For on-axis object. Relative Strehl ratio variations of 1.5-2.5% across entire IRIS field on account of multi-conjugate correction.
Airy ring size	J band: 21 mas H band: 28 mas K band: 37 mas	Diameter (FWHM)
Ensquared energy	J band: 0.35 - 0.57 H band: 0.50 - 0.66 K band: 0.62 - 0.72	Over 16.4 x 16.4 arcsec ² imager field. <i>Energy in box with diameter of PSF FWHM</i> Uncertainty originates from different conversions between WFE and EE.
Astrometric accuracy	Relative precision: 10 μ as Relative accuracy: 30 μ as Absolute accuracy: 2-4 mas	Relies on multiple visits to a particular field and a variety of reference fields.
Limiting magnitude (Imager)	J band: 27.8 H band: 27.3 K band: 26.9	Point source sensitivity. Five hour integration, S/N=100, 2 λ /D aperture. AB magnitude.
Limiting magnitude (Spectrograph)	J band: 25.8 H band: 26.0 K band: 25.2	Point source sensitivity for 4 mas pixel scale. Other scales are significantly more sensitive. Five hour integration, S/N=10, 2 λ /D aperture. AB magnitude.

